## A STUDY OF THE FRAGMENT DISPERSAL AND

 TRAJECTORY OF THE SAYH AL UHAYMIR 001 METEORITE SHOWER. A. V. Korochantsev ${ }^{1}$, D. A. Sadilenko ${ }^{1}$, M. A. Ivanova ${ }^{1}$, C. A. Lorentz ${ }^{1}$, and E.V. Zabalueva ${ }^{1}$, ${ }^{1}$ Vernadsky Institute of Geochemistry and Analytical Chemistry, Moscow 119991, Russia korochantzev@geokhi.ru.Sayh al Uhaymir 001 (SAUH 001) is one of Oman's largest known meteorite showers. We obtained hundreds of GPS find locations, analyzed the fragment distribution, and developed a model of the meteorite's trajectory.

SAUH 001 is a stony meteorite shower (L4/5) found March 16, 2000. More than 2670 samples weighing more than 450 kg have been collected. The collected samples may be a large portion of the total mass of the meteorite body, because the surface in this region is hard and it is likely that only a small portion of the fragments penetrated the surface. GPS locations were obtained for 748 found samples to within 0,001 minutes ( $\sim 2$ meters) (Fig. below). We assume that the fragments were not moved after falling.

The ground track of the meteorite was modeled as follows: (1) Determined center of mass (CM) of the fragment distribution. (2) Fitted a line, weighted for fragment mass, through the CM along the long axis of the distribution in both directions. Increasing fragment mass along track indicated a flight direction of $233^{\circ}$.

The fragments' mass distribution was used to model the meteorite trajectory. In the model, the distribution depends on the speed of entry into the atmosphere (v), angle to the surface ( $\alpha$ ), and breakup height $(\mathrm{H})$. We developed a computer program to calculate trajectories based on the distribution. For seven mass classes, we calculated the average mass and position along the ground track. We then varied $\mathrm{v}, \alpha$ and H to find solutions closest to the observed locations of the mass classes. To limit the solutions we used the relationship between v and H at the destruction air flux $\mathrm{A}=\left(\rho v^{2}\right)_{\max }$ :
$\mathrm{H}=-\mathrm{h} \cdot \ln \left(\mathrm{A} / \mathrm{v} \cdot \rho_{\mathrm{o}}\right)$, where $\rho_{\mathrm{o}}$ is the sea-level air density and h is the scale height. A was taken as $6.5 \cdot 10^{6}$, the average of several measurements for L chondrites [1]. The model yielded trajectory angle of $70-75^{\circ}$ and breakup heights of 40 km with a meteorite velocity of $27 \mathrm{~km} / \mathrm{s}, 30 \mathrm{~km}$ altitude with $\mathrm{v}=15 \mathrm{~km} / \mathrm{s}$, and 25 km with $\mathrm{v}=11 \mathrm{~km} / \mathrm{s}$. The meteorite may have broken up in several stages. For example, Pribram broke up gradually from 44 km down to 23 km altitude [2]. SAUH 001 samples show evidence of multi-stage fusion crust formation, suggesting multi-stage breakup.
Conclusion: Our model indicated the SAUH001 bolide travelled northeast to southwest, bearing $233^{\circ}$. Trajectory angle was $\sim 70^{\circ}$, with a probable velocity of $27 \mathrm{~km} / \mathrm{s}$. Our model gave a range of breakup altitudes similar to observations of Pribram. Breakup appears to have occurred in several stages.
References: [1] Tseplekha Z. (1961) Meteoritika 20, 178-184 (Russ.); [2] Tsvetkov V. I. (1985) Thesis, 120-122 (Russ.).


